Geophysical Research Abstracts, Vol. 10, EGU2008-A-04388, 2008 SRef-ID: 1607-7962/gra/EGU2008-A-04388 EGU General Assembly 2008 © Author(s) 2008



## Real-time Global MHD Simulations of Cassini T32 flyby

Y. J. Ma(1), C. T. Russell(1), A. F. Nagy(2), C. Bertucci(3), F. M. Neubauer(4), M.
K. Dougherty(3), T. E. Cravens(5), A. J. Coates(6), J-E. Wahlund(7), F. J. Crary(8)
(1)IGPP, UCLA, 6877 Slicher Hall, Los Angeles, CA 90025, United States
(yingjuan@igpp.ucla.edu; ctrussell@igpp.ucla.edu), (2)University of Michgan, 2455 Hayward
St., Ann Arbor, MI 48109, United States (anagy@umich.edu), (3)The Blackett Laboratory,
Imperial College, Space and Atmospheric Physics, London, SW7 2AZ, United Kingdom,
(4)University of Cologne, Rostocker Str 9, Cologne, 50374, Germany, (5)Univ. Kansas, 1251
Wescoe Hall, Lawrence, KS 66045, United States, (6)Mullard Space Science Laboratory,
University College London, MSSL-UCL Dorking Surrey, London, RH5 6NT, United
Kingdom, (7)Swedish Inst Space Physics, Uppsala Box 537, Uppsala, SE-75121, Sweden,
(8)Southwest research institution, 6220 Culebra Rd, San Antonio, TX 78228, United States

Cassini Spacecraft made a special pass of Titan on June 13, 2007. This flyby, referred to as T32 flyby, is the first occasion that Titan was observed to be outside Saturn's magnetopause (Bertucci et al., 2007). During the flyby, Titan encountered an abrupt change in the magnetic field. Other plasma parameters such as electron density and temperatures also changed dramatically during the pass (Coates et al., 2007). In this presentation, we study Titan's ionospheric responses to such a sudden change in the upstream plasma flow, using a sophisticated multi-species global MHD model. As a first attempt, we simplify the condition to a thin current sheet (represented by a tangential discontinuity) crossing, and simulate in real-time the interaction process of Titan's ionosphere with such a current sheet. Our simulation results show that the interaction causes magnetic reconnection at the upstream region and subsequently at the distant tail. There is a good agreement between the observations and the magnetic field predicted by our model, especially for the outbound segment. We also show results based on upstream conditions constrained by the Cassini plasma observations, as well as data-model comparisons.