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Turbulent heating and cooling of coronal loops

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We consider the coupling of both heating and cooling processes in coronal loops, including for the first time a feedback of the cooling on the heating. The heating is computed by a set of shell-models (see other contribution in this session) representing the non-linear interactions of MHD; they are coupled by Alfvén waves travelling along the loop from the footpoints, where the energy is introduced by motions of the photosphere. The shell-models display properties of anisotropic turbulence and intermittency. The heating computed by this model is then used as an input of a hydrodynamic model of a loop with thermal conduction and a self-consistent treatment of radiation (including non-equilibrium ion populations). The dynamic density profiles computed by this model are in turn used as an input for the heating model, providing thus a visible feedback on the heating.