Geophysical Research Abstracts, Vol. 8, 10208, 2006 SRef-ID: 1607-7962/gra/EGU06-A-10208 © European Geosciences Union 2006



Melt structures in comets

M. K. Wallis and N. C. Wickramasinghe Cardiff Centre for Astrobiology, Cardiff University (wallismk@cf.ac.uk / +44 29 2087 6436)

Since the three comets imaged at close quarters (comets Wild-2, Borrelly and Tempel-1) differ strongly in terrain types and overall topography, their history if not their origins were different. We see evidence of the three types of melting postulated for bodies containing substantial fractions of frozen H₂O: a) early central melting due to radionuclides, b) impact melts resulting from hypervelocity meteorites or cometesimals, and c) subcrustal melting due to solar heating of the dark (carbonaceous) outer surface. Though Tempel-1 currently has a similar orbit to the other two comets with q ~ 1.5 AU, its refrozen lakes imply that its dynamical history included a period of earth-crossing (q < 0.9 AU). Wild-2's lumpy nucleus with broad flat-bottomed 'craters' would be the remnant interior shell and partially recrystallised region resulting from central melting.