Geophysical Research Abstracts, Vol. 8, 01093, 2006 SRef-ID: 1607-7962/gra/EGU06-A-01093 © European Geosciences Union 2006



Variations of the magnetic and plasma kinetic energy densities

M. Leitner (1, 2), C. J. Farrugia (3, 4), V. A. Osherovich (5), J. Fainberg (5), H. K. Biernat (1, 2), R. Schwenn (6), K. W. Ogilvie (5)

(1) Space Research Institute, Austrian Academy of Sciences, A–8042 Graz, Austria,(2) Institute for Geophysics, Astrophysics, and Meteorology, University of Graz, A–8010 Graz, Austria, (3) Institute of the Study of Earth, Oceans and Space, University of New Hampshire, Durham, NH 03824, USA, (4) Physics Department, University of New Hampshire, Durham, NH 03824, USA, (5) Laboratory for Extraterrestrial Physics, NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA, (6) Max-Planck-Institute für Sonnensystemforschung, Katlenburg Lindau, D–37191, Germany. (martin.leitner@oeaw.ac.at / phone: +43 316 4120 633)

A quantity called the quasi-invariant (QI), defined as the ratio of the magnetic energy density to the plasma kinetic energy density, has been found to be a good proxy for solar activity. This quantity was found to correlate strongly with the sunspot number (cc > 0.9), a correlation which holds not only for observations at 1 AU but also at Venus at 0.72 AU. However, no extensive observations on the QI index were reported in the inner heliosphere. Using the entire Helios data set we examine QI as a function of radial distance in the range 0.3–1 AU. A further new aspect is that we factor in the contribution of He++ ions to this parameter. Finally, the difference between QI values in the ambient solar wind, and those obtained during magnetic cloud events are explored and discussed.